

DAILY RADIO FREQUENCY OBSERVATIONS OF SELECTED OBJECTS

B.J. Geldzahler^{1,3}, K.J. Johnston¹, J.H. Spencer¹,
E.B. Waltman¹, W.J. Klepczynski², P.E. Angerhofer²,
D.R. Florkowski², D.D. McCarthy², and D.N. Matsakis²
¹E. O. Hulburt Center for Space Research, Naval Research
Laboratory, Washington, D. C. 20375
²U. S. Naval Observatory, Washington, D. C. 20390
³NRC/NRL Cooperative Research Associate

In Table I we present the list of 38 celestial objects that have been observed since January 1978 at 2.7 and 8.1 GHz with the Green Bank interferometer. The sources fall naturally into three categories: radio stars, possibly Galactic sources, and extragalactic sources. SS433, Cyg X-3, and each extragalactic source is measured several times per day while the other sources are measured once every three days. Reports on the entire program will be found in Geldzahler *et al.* (1983a), and on specific sources: SS433--Johnston *et al.* (1983a), BL Lac--Johnston *et al.* (1983b), Cyg X-3--Geldzahler *et al.* (1983b) and elsewhere in this volume), and CTA 26--Spencer *et al.* (1983).

We have defined for the variable sources a "rapidity index" which gives the number of maxima/year. This index includes major outbursts as well as "flickering". We also show in Table I the value of k ($-d(\log S_{\text{max}})/d(\log \lambda)$). The values of this index fall into three groups: $k > 0$, $k \sim 0$, and $k \sim -0.4$. A uniform source that is initially optically thick and whose energy losses occur primarily through adiabatic expansion should yield $k = -1$ (c.f. van der Laan 1966). We find that $k > 0$ when we have an optically thin object such as SS433 or Cyg X-3 during outburst. To make the value of $k \sim -0.4$ more agreeable with the standard model, we suggest the uniformity should be replaced by a variable opacity throughout the source. Finally $k \sim 0$ in those sources, such as the "quiescent" Cyg X-3, where repeated, rapid flickering has stretched and weakened the magnetic field in the immediate vicinity of the source.

REFERENCES

- Geldzahler, B.J. *et al.*: 1983a, to be submitted to A. J.
Geldzahler, B.J. *et al.*: 1983b, Ap. J. (Letters) in press.
Johnston, K.J. *et al.*: 1983a, A. J. in press.
Johnston, K.J. *et al.*: 1983b, Ap. J. (Letters) in press.
Spencer, J.H. *et al.*: to be submitted to Ap. J.
van der Laan, H.: 1966, Nature 211, 1131.

179

Table I. List of Program Sources

| Source | ID | Observing Interval ^a | Redshift | b | c | d | Other Names |
|---|----|---------------------------------|----------|-------|------------|------------|--|
| | | | | RI | k | | |
| a. Radio Stars | | | | | | | |
| 0236+610 | | 5001-5007 | | I | | | LSI 61 +303 |
| 0323+285 | | 4976-5006 | | -- | | | UX A-1 |
| 0334+004 | | 4971-5010 | | 28.1: | | | HR1099 |
| 1617-155 | | 4977-5517 | | -- | | | Sco X-1 |
| 1909+048 | | 4067-5017 | | 5.7 | +0.53±0.03 | | SS433 |
| 1956+350 | | 4976-5515 | | -- | | | Cyg X-1 |
| 2050+407 | | 4971-5517 | | -- | -0.04±0.5 | | Cyg X-3 |
| 2259+585 | | 4974-5129 | | -- | | | GF2259+586 |
| b. Possibly Galactic Objects | | | | | | | |
| 0125+628 | | 4976-5517 | | 4.0 | | | G127.11+0.54 |
| 2013+370 | | 4983-5517 | | 3.7 | | | G74.89+1.22 |
| c. Extragalactic Objects^a | | | | | | | |
| 0224+671 | Q | 3942-5517 | | | 3.6 | | |
| 0235+164 | BL | 4971-5148 | | | 4.1 | -0.42±0.05 | DD160 |
| 0237-234 | Q | 3941-5148 | 2.223 | | 1.5 | | DD-263, PHL8462 |
| 0316+413 | C | 3942-4066 | | | I | | 3C84 |
| 0336-019 | Q | 3942-5148 | 0.852 | | 6.0 | -0.01±0.03 | CTA26 |
| 0355+508 | EF | 3941-5148 | | | 3.1 | | |
| 0402-382 | Q | 3941-4066 | 1.417 | | 8.7: | | |
| 0727-115 | | 5060-5148 | | | 24.9: | -0.09±0.02 | |
| 0742+103 | EF | 3942-4066, 5010-5059 | | | I | | |
| 0851+203 | BL | 3942-5148 | 0.306: | | 3.9 | -0.48±0.01 | OJ287 |
| 0923+392 | Q | 3942-4066 | 0.699 | | I | | C39.25, DA267, DK340 |
| 0964+658 | BL | 3942-5517 | | | 2.1 | -0.04±0.02 | |
| 1226+023 | Q | 3943-5517 | 0.158 | | I | | 3C273, 4C02.32, NRA0400, DN044, DA324 |
| 1245-197 | Q | 3942-5517 | | | 0.0 | | |
| 1328+254 | Q | 3942-5517 | 1.055 | | 0.0 | | 3C287, 4C25.43, NRA0424, OP247, DA345 |
| 1328+307 | Q | 3942-5517 | 0.849 | | 0.0 | | 3C286, 4C30.26, NRA0425, OP348, DA346, CTA60 |
| 1502+106 | Q | 3942-5517 | 1.833 | | 2.7 | +0.08±0.03 | DR103 |
| 1519-273 | Q | 3943-4067 | | | 11.8: | | |
| 1641+399 | Q | 3941-5517 | 0.595 | | 0.6 | | 3C345, 4C39.48, NRA0513, OS368, DA420 |
| 1749+701 | BL | 3943-5517 | | | 2.4 | -0.03±0.04 | W1 |
| 1901+319 | Q | 3943-4066 | | | I | | 3C395 |
| 2021+614 | Q | 5010-5059 | | | I | | |
| 2037+511 | Q | 3943-4066 | 1.686 | | I | | 3C418, 4C51.12, NRA0636 |
| 2048+312 | Q | 4970-5517 | 3.18 | 22.3 | | | CL4 |
| 2134+004 | Q | 4011-5517 | 1.936 | | I | | PHL61, DA553, DX057 |
| 2200+420 | BL | 3942-5517 | 0.0688 | 5.7 | -0.30±0.09 | | BL Lac |
| 2251+158 | Q | 3942-5517 | 0.859 | 3.9 | -0.12±0.01 | | 3C454, 3.4C15.76, OY185, NRA0701, DA506 |
| 2345-167 | Q | 3942-4066 | 0.600 | 20.6: | | | DZ-176 |

a As of 1 July 1983

b Emission line redshifts taken from Hewitt and Burbidge (1980)

c Rapidity Index: the number of maxima per year; I = variations exist by the value of RI cannot be determined with reliability, -- = no obvious variation, : = value is uncertain due to short time base

d $k = (d \log \text{peak flux density}) / (d \log \text{frequency})$

e In Cyg X-3, $k \approx 0.5$ during optically thin outbursts and $k \approx -0.4$ during optically thick outbursts