## TOWARDS A COMPLETE LIBRARY OF STELLAR SPECTRA FOR EVOLUTIONARY SYNTHESIS

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The population and evolutionary synthesis of the integrated light of clusters and galaxies requires a good knowledge of the underlying stellar spectra. Libraries of observed stellar spectra can be used, but they have several disadvantages, e.g. uneven sampling — which causes problems in the integration phase. Furthermore, no comprehensive library of low—or highmetallicity stars does exist, which would be required to model chemical evolution. Libraries of synthetic spectra could — and should — solve these problems. Using Kurucz and more recent libraries dedicated to cool stars, we constituted a library covering the whole range of  $T_{\rm eff}$ : 50 000 – 2500 K.

The comparaison between empirical and model computed temperature-colors relations shows differences that are too big to allow to use the cool synthetic spectra as such. We therefore developed a method to modify the synthetic spectra, in order to allow them to fit the observed temperature-colors relations.

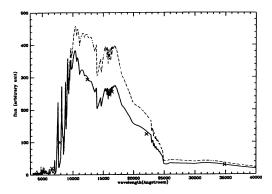


Figure 1. Original and modified spectrum at  $T_{\rm eff} = 3126$ K. The crosses show the empirical fluxes in the UBVRIJHKL bands at this temperature.