

THE PRESENT DATA SITUATION FOR STARS IN OPEN CLUSTERS. II.

J.-C. Mermilliod
Institut d'Astronomie de l'Université de Lausanne
et Observatoire de Genève

The present state of astrometric, spectroscopic and photometric data for stars in 63 open clusters nearer to the Sun than 750 pc is summarized. Table I reports the limits of the available data in terms of the apparent V magnitude, in the same way as the previous version (Mermilliod 1977). Information on the apparent V magnitude of the brightest main sequence star (V_{br}) and on the number of stars brighter than $V=10$ (n^*) has been included. In addition, the bluest spectral type on the main sequence has been used as an age estimator (TS).

Distances (d) and colour excesses ($EB-V$) result from new determinations by the author for one half of the clusters. Otherwise, they have been taken from the catalogues of Becker and Fenkart (1971) and Fenkart and Binggeli (1979).

Columns 6 to 8 contain the V magnitudes of the faintest member star reached in three spectroscopic domains, i.e. the spectral classification (MK) with indication of the corresponding type (latest), the projected rotational velocity ($V\sin i$) together with the intrinsic colour index of the reddest star reached ($(B-V)_0$), and the radial velocity (VR).

Details about the limits in the proper motion data are to be found under the heading (pm).

The last three columns give the V magnitudes of the faintest star measured in the UBV system, photoelectrically (UBV) and photographically (pg), in the Geneva (7-col) and Strömrgren (4-col.) systems, both with an indication of the number of measured stars (no.).

Cluster	α 1950	λ II	d	TS	Vbr	MK	$V \sin i$	VR	UBV	7-col.	4-col.
	δ 1950	b III									
UMa Cl.	14 ^h 40	110 ^o .2	22	A1	1.8	8.3	6.8	8.3	8.3	8.3	8.3
	+69 ^o 47	+44 ^o .8	.00		14	K3V	0.6	8.3	-	14	14
Hyades	4 17	179.1	43	A2	3.5	12.4	7.5	9.5	15.3	12.3	9.2
	+15	-23.9	.00		150	MOV	0.6	18.3	18.3	170	92
Coma Ber	12 23	221.1	89	A2	5.0	11.2	9.7	9.7	15.6	11.7	10.6
	+26 23	+84.1	.00		37	G9V	0.7	13.	-	65	51
Pleiades	3 44	166.6	130	B6	2.9	15.3	10.4	9.2	18.7	11.6	10.3
	+23 58	-23.5	.04		75	M2.5V	0.6	16.	18.5	177	91
IC 2391	8 39	270.4	150	B3	3.6	9.6	8.6	9.9	12.4	10.8	10.2
	-52 53	-6.9	.00		25	F5V	0.2	-	-	39	44
IC 2602	10 41	289.6	150	B3	4.8	10.3	7.8	-	10.7	10.9	10.3
	-64 08	-4.9	.04		32	G2IV	0.1	-	-	43	34
α Per	3 20	147.0	170	B5	4.2	11.0	9.7	9.7	11.7	11.9	11.7
	+49	-6.0	.08		80	G1V	0.4	12.0	-	104	97
Praesepe	8 37	205.5	174	A2	6.3	9.9	10.3	10.3	13.9	12.6	10.8
	+20 10	+32.5	.00		70	F8V	0.5	16.	-	130	99
NGC 2451	6 44	252.4	200	B3	4.8	9.4	-	-	13.0	9.3	6.5
	-37 51	-6.7	.00		25	F0V	-	12.	-	28	11
Zeta SC1	0 02	14.9	240	B8	7.0	8.9	-	8.9	10.2	12.5	12.5
	-30 18	-79.2	.00		18	A3V	-	-	-	57	48
NGC 6475	17 51	355.9	260	B9	5.9	9.5	8.5	8.5	12.3	11.1	11.6
	-34 48	-4.5	.06		40	A5V	0.1	-	11.4	105	119
NGC 7092	21 30	92.5	305	A0	6.5	10.1	9.1	9.1	10.9	11.2	-
	+48 13	-2.3	.01		23	F5V	0.1	15.	-	32	-
IC 348	3 41	160.4	316	B1	3.8	11.8	-	-	13.0	12.8	-
	+32 08	-17.7	Var		4	A2V	-	13.4	-	11	-
NGC 2232	6 24	214.2	320	B2	5.0	9.6	9.6	-	11.4	-	-
	-4 43	-8.1	.02		24	A1V	0.1	-	-	-	-
Cr 132b	7 12	243.3	330 (B6)		7.5	-	-	-	9.5	-	-
	-31 05	-9.2	.03		8	-	-	-	-	-	-
IC 4665	17 44	30.6	350	B3	6.8	10.6	10.6	10.6	11.6	11.1	11.4
	+5 44	+17.1	.18		20	A5IV	0.2	14.0	13.7	28	45
Rup 98	11 55	297.2	350 (A1)		8.9	-	-	-	11.8	-	-
	-64 12	-2.2	.17		6	-	-	-	-	-	-
NGC 6633	18 25	36.1	360	A2	8.4	9.7	-	-	15.1	10.4	10.9
	+6 32	+8.3	.11		25	A5V	-	15.5	14.0	28	38

Cluster	α 1950 δ 1950	ℓ II b II	d $EB-V$	TS	Vbr n*	MK latest	$V \sin i$ $(B-V)_0$	VR pm	UBV pg	7-col. no.	4-col. no.
Cr 140	7 ^h 22	245 ^o 0	378	B2	5.3	6.8	-	-	12.0	-	-
	-31 ^o 56	-8 ^o 0	.05		18	B5V	-	-	-	-	-
NGC 752	1 55	137.2	380	F2	9.2	11.7	10.1	12.5	13.9	13.9	11.7
	+37 26	-23.4	.03		10	F5V	0.4	14.9	15.5	106	37
NGC 2547	8 09	264.5	380	B3	6.5	10.5	-	-	-	11.4	-
	-49 07	-8.5	.05		26	A7V	-	11.5	-	57	-
Orion C1	5 33	206.7	400	O9.5	2.7	13.3	10.5	11.7	15.3	11.9	12.0
	-5 25	-20.7	.06		55	K1IV	0.5	15.	18.0	73	125
NGC 2516	7 57	274.0	400	B8.5	5.6	10.8	9.4	8.6	12.3	12.2	12.4
	-60 46	-15.9	.11		62	A8	0.1	-	-	102	134
IC 4756	18 37	36.4	400	(A5)	9.5	-	-	-	11.0	-	-
	+5 26	+5.3	.19		20	-	-	14.5	14.1	-	-
NGC 3228	10 20	280.7	415	B6	7.9	10.2	-	-	14.5	-	-
	-51 28	+4.6	.00		11	A1V	-	-	-	-	-
NGC 1662	4 46	187.7	420	A1	9.0	10.7	-	-	15.1	-	-
	+10 51	-21.1	.29		8	A3V	-	-	15.4	-	-
Tr 10	8 46	262.8	440	B2	6.4	9.7	-	-	11.0	-	-
	-42 18	+0.6	.02		20	A6V	-	-	-	-	-
NGC 3532	11 04	289.6	445	A0	7.4	10.9	-	-	11.9	11.0	-
	-58 24	+1.5	.02		50	A7V	-	11.5	13.5	102	-
NGC 2422	7 43	231.0	456	B6	6.2	10.7	9.4	10.4	13.8	10.8	-
	-14 23	+3.1	.07		20	A2V	-0.1	12.0	14.2	22	-
NGC 6405	17 37	355.6	460	(B5)	6.7	-	-	-	14.5	11.4	-
	-32 11	-0.7	.15		28	-	-	-	14.7	52	-
NGC 1039	2 39	143.6	470	B9.5	8.0	9.6	9.6	9.6	12.2	12.0	-
	+42 34	-15.6	.09		22	A0V	0.0	13.8	-	43	-
NGC 2281	6 46	175.0	496	A1	8.6	-	-	-	14.9	12.9	-
	+41 07	+17.1	.07		5	-	-	12.8	13.1	58	-
NGC 6281	17 04	347.8	525	(B9)	8.3	-	-	-	13.5	10.8	-
	-37 49	+2.0	.14		12	-	-	-	-	16	-
NGC 1647	4 43	180.4	550	B9	8.6	10.7	-	-	16.1	-	-
	+18 59	-16.8	.39		10	A0IV	-	-	16.3	-	-
Pi 4	8 33	262.7	550	(B6)	7.5	-	-	-	11.9	-	-
	-44 06	-2.4	.03		4	-	-	-	-	-	-
Cr 132a	7 12	243.3	560	B2	6.3	-	-	-	10.6	-	-
	-31 05	-9.2	.02		10	-	-	-	-	-	-

Cluster	α 1950 δ 1950	λ II b II	d $EB-V$	TS	Vbr n*	MK latest	$V \sin i$ $(B-V)_0$	VR pm	UBV PG	7-col. no.	4-col. no.
Ma 6	2 ^h 26	134. ⁰ 7	560	(B1)	8.5	-	-	-	-	-	-
	+60 ^o 26	+0. ⁰ 0	.64		4	-	-	-	16.3	-	-
NGC 6451	17 47	359.5	565	(F)	RGU	-	-	-	-	-	-
	-30 12	-1.6	.09		-	-	-	-	-	-	-
NGC 6124	16 22	340.8	580	(B9)	8.7	-	-	-	-	-	-
	-40 35	+6.0	.68		13	-	-	-	14.0	-	-
NGC 2527	8 03	246.1	590	A2	9.4	-	-	-	13.7	-	-
	-28 01	+1.9	.08		5	-	-	-	16.6	-	-
NGC 6167	16 30	335.3	590	(B6)	9.0	-	-	-	15.6	-	-
	-49 30	-1.3	.89		9	-	-	-	15.7	-	-
NGC 1342	3 28	155.0	600	A1	8.7	10.7	-	-	15.0	-	-
	+37 10	-15.4	.30		2	A2V	-	-	15.1	-	-
NGC 6716	18 52	15.4	600	B8	8.3	-	-	-	13.8	-	-
	-19 57	+9.6	.13		5	-	-	-	13.9	-	-
IC 4725	18 29	13.6	600	B5	7.9	10.9	-	10.3	16.3	-	11.3
	-19 17	-4.5	Var		25	A1V	-	-	16.0	-	10
Ba 19	13 25	307.4	600	(B9)	RGU	-	-	-	-	-	-
	-61 00	+1.3			-	-	-	-	-	-	-
Tr 2	2 34	137.4	615	B7	8.4	10.8	-	-	15.3	-	-
	+55 46	-3.9	.33		10	AOV	-	-	15.7	-	-
NGC 225	0 40	122.0	630	B8	9.3	10.9	-	-	15.4	-	-
	+61 31	-1.1	.29		4	AOV	-	-	15.9	-	-
NGC 5367	13 55	316.5	630	(B4)	9.1	-	-	-	13.3	-	-
	-39 44	+21.1	Var		2	-	-	-	-	-	-
NGC 7063	21 22	83.1	630	B8	8.9	10.8	-	-	15.4	-	-
	+36 17	-9.9	.08		4	A1V	-	-	15.9	-	-
NGC 2548	8 11	227.9	640	(A1)	8.8	-	-	-	11.5	-	-
	-5 39	+15.3	.00:		12	-	-	-	-	-	-
NGC 3114	10 01	283.3	640	B8	7.7	10.4	-	-	11.7	10.2	-
	-59 52	-3.8	.10		26	B9IV	-	-	12.0	26	-
Cr 463	1 44	127.4	650	(B6)	8.7	-	-	-	12.8	-	-
	+71 42	+9.6	Var		14	-	-	-	-	-	-
NGC 2287	6 45	231.1	680	B9	7.8	10.8	-	11.1	14.0	10.7	11.8
	-20 41	-10.2	.01		25	A2V	-	-	14.4	45	16
NGC 6494	17 54	9.8	690	A0	9.3	10.0	-	-	15.6	-	-
	-19 01	+2.9	.38		17	A5V	-	-	14.1	-	-

Cluster	α 1950 δ 1950	λ II b II	d <i>EB-V</i>	TS	Vbr n*	MK latest	$V_{\text{sin}i}$ $(B-V)_0$	VR pm	UBV pg	7-col. no.	4-col. no.
NGC 5662	14 ^h 32	316 ^o 09	700	B7	8.3	11.0	-	-	15.3	-	-
	-50 ^m 20	+3 ^s 05	.32		10	B9V	-	-	15.2	-	-
NGC 6134	16 24	334.4	700 (B8)	11.1		-	-	-	14.8	-	-
	-49 02	-0.2	.45		-	-	-	-	14.9	-	-
NGC 5460	14 04	315.8	720 (B8)	8.0		-	-	-	13.1	10.7	-
	-48 05	+12.7	.13		11	-	-	-	-	11	-
NGC 5823	15 02	321.2	720 (F?)	10.9		-	-	-	14.2	-	-
	-55 14	+2.6	.27		-	-	-	-	14.0	-	-
Rup 108	13 29	308.2	730 (B9)	9.3		-	-	-	-	-	-
	-58 14	+4.0	.18		3	-	-	-	-	-	-
NGC 5822	15 01	321.7	740 (A5)	10.2		-	-	-	13.7	-	-
	-54 09	+3.6	.14		-	-	-	-	13.9	-	-
NGC 7160	21 52	104.0	740 B1	6.7		13.1	9.3	13.1	14.5	12.9	-
	+62 22	+6.5	.36		7	A7V	-0.2	-	14.9	24	-
NGC 2925	9 32	276.0	745 (B8)	8.4		-	-	-	-	-	-
	-53 13	-1.2	.06		3	-	-	-	-	-	-
Rup 46	8 00	283.4	750 (F)	11.2		-	-	-	14.1	-	-
	-19 20	+5.9	.07		-	-	-	-	-	-	-

Information on proper motions, MK classification and *UBV* photometry has been derived from the catalogues of Mermilliod (1976a, b) and from the third edition in preparation. Data concerning the radial velocities have been extracted from the compilation by Mermilliod (1979) and its continuation, and for the $V_{\text{sin}i}$ from an unpublished file by the author.

Information relative to the Geneva seven-colour (*U, B, V, B1, B2, V1, G*) and Strömrgren *wby* photometric systems has been kindly communicated respectively by F. Rufener and M. Mermilliod in advance of publication.

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