

Design and field testing of the Fish-Eye lens for optical atmospheric observations

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Abstract. The Fish-Eye lens MAO-08 is intended for observations of weak extended objects (aurora, twilight and dawn phenomena, stratospheric clouds, etc.) in narrow spectral bands with variable passband filters VARISPEC and in white light. It is valuable for astronomical observation for the target – background problem when we need to estimate the spectral transparency of the atmosphere. Besides having high power this lens can be used for meteor observations. It has been tested during the winter field conditions.

Keywords. Fish-Eye lens; Interference Filter; Auroral Observations, Hyaline Observations.

1. Introduction

The Fish-Eye lens is most often used for observations of atmospheric phenomena such as aurora and nightglow (Elvey & Stoffregen(1957), Lebedinsky (1961), Sandahl *et al.* (2008)). For observations of broad band, commercial fish-eye lenses can be used coupled to CCD detector. However, when it comes to obtaining narrow band images of auroral emissions lines with high temporal resolution, the task of researchers is much more complicated. The filters require to be collimated light in order to obtain optimum bandpass. A standard technique is to use an telecentric lens system (Mende *et al.* (1977), Sandahl *et al.* (2008)). The MAO-08 lens does not use this system and based on calculations as a whole.

2. Main Parameters

The ultra wide-angle lens MAO-08 can be used with interference filters, variable spectral bandwidth filters (VARISPEC), or without any filters (white light). Figure 1 shows the optical diagram of the lens. The first part converts the 180° field of view (FOV) to a narrow beam of 6°, where interference filters may be inserted. The back part focuses the collimated light onto the detector plane. The design is compact with unique optical characteristics. The basic technical data:

- FOV: 180 degrees;
- F/value: 0.82;
- Spectral range: 430 – 750 nm;

- Resolution: center – 100 l/mm; edge – 70 l/mm;
- Dimensions (with covers): length 235 mm and diameter 103 mm

3. Field testing

Lens tests in field conditions have confirmed high quality of images. (Fig. 2a) is image of the ray auroral arc, obtained at Barentsburg in white light. There are a plenty of constellations and stars in this figure and even the Milky Way is seen. (Fig. 2b) shows the weak auroral folds in the emission at 557.7 nm, obtained with narrow-band interference filter ($\Delta\lambda \sim 2$ nm).

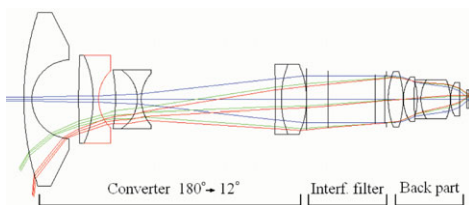


Figure 1. Layout of the MAO-08 lens.

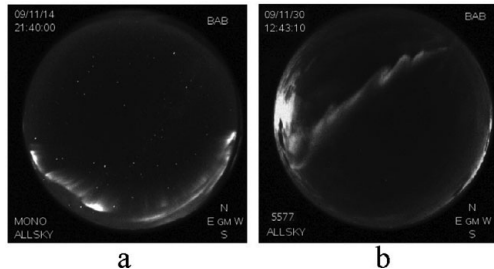


Figure 2. All-sky images of aurora: a) in white light, b) at $\lambda \sim 557.7$ nm.

A first season of observations was spent during the period November 2011 – April 2012 in the Kjell Henriksen Observatory (KHO) in Svalbard located at the archipelago Svalbard 1000 km north of mainland Norway (78°N, 16°E), and Barentsburg research Station, PGI (78.093°N, 14.208°E).

4. Conclusion

The MAO-08 lens has a large aperture, which exceeds this parameter of the other lenses for auroral research. It allows filters to be inserted into the lens. Pointed out is the possibility of obtaining simultaneously two images in two emissions of the same aurora by the single CCD camera.

5. Acknowledgements

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