

HIGH RESOLUTION SPECTRAL STUDIES OF SOME DIATOMIC MOLECULES OF ASTROPHYSICAL INTEREST

T.K. Balasubramanian, G.L. Bhale, Sheila Gopal, G. Krishnamurthy, G. Lakshminarayana, M.D. Saksena, P. Saraswathy, B.J. Shetty and Mahavir Singh
Spectroscopy Division, Bhabha Atomic Research Centre,
Bombay-400 085, India.

ABSTRACT. Laboratory studies have been made on molecules of astrophysical interest such as AlO, CO, CrO, SiS, NH⁺ and OH. Vibrational and rotational constants have been determined more accurately in the various electronic states.

Laboratory studies of molecular spectra are essential for identifying molecules in stellar atmospheres, in interstellar medium and in the planetary atmosphere. The interplay between molecular spectroscopy and astrophysics has often been emphasized (1,2). In our laboratory we have carried out high resolution spectral studies of molecules of astrophysical interest such as AlO, CrO, SiS, NH⁺, CO and OH.

The presence of AlO was detected in cool stars and sun-spots through the identification of B²Σ⁺ - X²Σ⁺ transition in blue-green region. The intensity of the 0-0 band in Mira stars is highly variable and at times the band appears in emission. It is suggested that if another electronic transition such as A²Π_i - X²Σ⁺ is observed in the stellar spectra, along with the blue-green¹ transition, it would be possible to understand the temperature changes in the stellar atmosphere and the excitation mechanism. Luck and Lambert (3) have calculated the 1-0 band of the A²Π_i - X²Σ⁺ transition and looked for it in the stellar spectra. They attributed their negative results to the inaccurate rotational constants of the A²Π_i state. We have made an extensive study of four electronic transitions viz. E²Δ_i - A²Π_i, C²Π_r - X²Σ⁺, C²Π_r - A²Π_i and D²Σ⁺ - A²Π_i of AlO (4), and obtained accurate rotational constants for the A²Π_i state which would help in the positive identification of the A²Π_i - X²Σ⁺ transition in stellar atmosphere. Our studies have also resulted in an unambiguous analysis of the rotational structure of v = 0,1 and 2 of the C²Π_r state. Prior to our studies no reliable rotational constants existed for the C²Π_r state (5).

Similar studies have been done on the high resolution spectra of several molecules like CrO (6), NH⁺ (7), SiS (8,9,10). Theoretical

expressions for the rotational line intensities are derived for a ${}^2\Sigma - {}^2\Pi$ transition and applied to OH molecule (11). From isotope shift studies of CO the Kaplan bands are shown to belong to the $v' = 2$ progression of the 3rd positive system $b{}^3\Sigma^+ - a{}^3\Pi$, thus confirming the speculation of Schmid and Gero (12,13,14).

REFERENCES

1. G. Herzberg, J. Opt. Soc. Amer. 55, 229 (1965).
2. G. Herzberg, High-lights of Astronomy, ed. Patrick A. Wayman, 5, 3 (1980).
3. R. Earle Luck and David L. Lambert, Publi. Astron. Soc. Pacific 86, 276 (1974).
4. M. Singh and M.D. Saksena, Can. J. Phys. 59, 955 (1981), 60, 1730 (1982), 61, 1347 (1983), 63, (1985) in press.
5. K.P. Huber and G. Herzberg, Molecular Spectra and molecular structure - constants of diatomic molecules Vol. IV, Van.Nostrad Reinhold Company, New York, N.Y. (1979).
6. W.H. Hocking, A.J. Merer, D.J. Milton, W.E. Jones and G. Krishnamurthy, Can. J. Phys. 58, 516 (1980).
7. Krishnamurthy and M. Saraswathy, Pramana, 6, 235 (1976).
8. Sheila Gopal, G. Lakshminarayana and N.A. Narasimham, J. Phys. B 13, 3781 (1980).
9. G. Krishnamurthy, Sheila Gopal, P. Saraswathy and G. Lakshminarayana Can. J. Phys., 61, 714 (1983).
10. G. Lakshminarayana, B.J. Shetty and Sheila Gopal, J. Molec. Spect. 112, 1 (1985).
11. T.K. Balasubramanian, A.S.P. Rao, R.D¹Souza and N.A. Narasimham, Acta Physica Hungarica 55, 27 (1984)
12. G.L. Bhale (Private communication)
13. J. Kaplan, Phys. Rev. 35, 1298 (1930).
14. R. Schmid and L. Gero, Z. Physik, 99, 281 (1936).