

Halo K-Giant Stars from LAMOST: Kinematics and Galactic Mass Estimate

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Abstract. We analyze line-of-sight velocities of over 3000 halo K-giant stars from the second data release of the spectral survey LAMOST (Zhao *et al.* 2012). We find a nearly constant velocity dispersion profile, with no large dips or peaks, in a Galactocentric radial range of 10 – 30 kpc, in accord with earlier analyses (Battaglia *et al.* 2005, 2006; Xue *et al.* 2008, 2014) (see Fig. 1). Previous studies of halo star radial velocity dispersions in a reference frame centered on the Galactic Center have detected dips within this radial range (Sommer-Larsen *et al.* 1994; Kafle *et al.* 2012, 2014). We use the stars to make estimates of the enclosed mass out to 40 kpc from the Galactic Center using the method of Evans *et al.* (2011). Tens of thousands of such stars are expected to become available to this analysis by the end of the five-year survey.

Keywords. Galaxy: halo — stars: individual (K giants) — stars: kinematics and dynamics

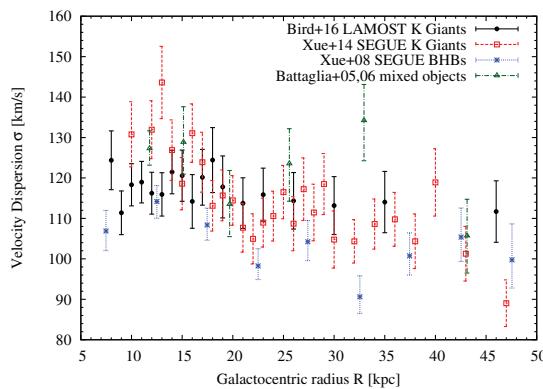


Figure 1. Comparison of line-of-sight halo velocity dispersions from Bird *et al.* (2016) LAMOST K giants, Xue *et al.* (2014) SEGUE K giants, Xue *et al.* (2008) SEGUE blue horizontal branch stars (BHBS), and Battaglia *et al.* (2005, 2006) mixed objects as a function of Galactocentric radius in spherical coordinates. The profiles follow a flattened distribution.

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