

Detecting Low-Order CO Emission from $z \gtrsim 4$ Quasar Host Galaxies

D. A. Riechers¹, F. Walter¹, C. L. Carilli², K. K. Knudsen¹,
K. Y. Lo³, D. J. Benford⁴, J. G. Staguhn⁴, T. R. Hunter⁵,
F. Bertoldi⁶, C. Henkel⁷, K. M. Menten⁷, A. Weiss⁷,
M. S. Yun⁸, and N. Z. Scoville⁹

¹Max-Planck-Institut für Astronomie, Heidelberg, Germany

²National Radio Astronomy Observatory, Socorro, NM, USA

³National Radio Astronomy Observatory, Charlottesville, VA, USA

⁴Lab. for Observational Cosmology, NASA Goddard Space Flight Center, Greenbelt, MD, USA

⁵Harvard-Smithsonian Center for Astrophysics, Cambridge, MA, USA

⁶Argelander-Institut für Astronomie, Bonn, Germany

⁷Max-Planck-Institut für Radioastronomie, Bonn, Germany

⁸Department of Astronomy, University of Massachusetts, Amherst, MA, USA

⁹Astronomy Department, California Institute of Technology, Pasadena, CA, USA

Abstract. Molecular gas has now been detected in 15 $z > 2$ QSOs. These detections are commonly obtained by observing high- J CO transitions due to their relatively high peak fluxes and observing frequencies in the millimeter atmospheric windows. However, only observations of the CO ground-state transition, CO(1–0), have the potential to trace the molecular gas at lower excitations, which may give a better estimate of the total molecular gas mass of high- z QSOs. Here we present first $z > 4$ CO(1–0) observations obtained with the NRAO Green Bank Telescope and the MPIFR Effelsberg telescope (Riechers *et al.* 2006). With these two 100m telescopes, we detect the CO(1–0) transition in the high-redshift QSOs BR 1202-0725 ($z = 4.7$), PSS J2322+1944 ($z = 4.1$), and APM 08279+5255 ($z = 3.9$). We find that the CO/FIR luminosity ratios of these high- z sources follow the same trend as seen for low- z galaxies. Utilizing large velocity gradient (LVG) models based on previous results for higher- J CO transitions, we derive that all CO emission can be described by a single gas component and that all molecular gas appears to be concentrated in a compact nuclear region. We thus find no evidence for luminous, extended CO(1–0) components in the molecular gas reservoirs around our target quasars.

Keywords. cosmology: observations, galaxies: ISM, galaxies: high-redshift, galaxies: formation, galaxies: starburst, galaxies: active

Acknowledgements

D. R. acknowledges support from the Deutsche Forschungsgemeinschaft (DFG) Priority Programme 1177. C. C. acknowledges support from the Max-Planck-Gesellschaft and the Alexander von Humboldt-Stiftung through the Max-Planck-Forschungspreis.

Reference

Riechers, D. A., Walter, F., Carilli, C. L., *et al.* 2006, *ApJ* in press (astro-ph/0606422)